## **Glassy Dynamics in Chiral Fluids**

Vincent E. Debets<sup>0</sup>, Hartmut Löwen<sup>0</sup>, and Liesbeth M. C. Janssen<sup>0</sup>

<sup>1</sup>Department of Applied Physics, Eindhoven University of Technology, P.O. Box 513, 5600 MB Eindhoven, Netherlands <sup>2</sup>Institut für Theoretische Physik II: Weiche Materie, Heinrich-Heine-Universität Düsseldorf, D-40225 Düsseldorf, Germany

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Chiral active matter is enjoying a rapid increase of interest, spurred by the rich variety of asymmetries that can be attained in, e.g., the shape or self-propulsion mechanism of active particles. Though this has already led to the observance of so-called chiral crystals, active chiral glasses remain largely unexplored. A possible reason for this could be the naive expectation that interactions dominate the glassy dynamics and the details of the active motion become increasingly less relevant. Here, we show that quite the opposite is true by studying the glassy dynamics of interacting chiral active Brownian particles. We demonstrate that when our chiral fluid is pushed to glassy conditions, it exhibits highly nontrivial dynamics, especially compared to a standard linear active fluid such as common active Brownian particles. Despite the added complexity, we are still able to present a full rationalization for all identified dynamical regimes. Most notably, we introduce a new "hammering" mechanism, unique to rapidly spinning particles in high-density conditions, that can fluidize a chiral active solid.

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*Introduction.*—Inspired by its omnipresence in biology, as well as its growing relevance in condensed matter and materials science, active matter has proven to be one of the prevailing subjects in biological and soft matter physics [1–3]. Active or self-propelled particle systems are intrinsically far from equilibrium, giving rise to a myriad of surprising features that are inaccessible to conventional passive matter. Well-known examples include motility induced phase separation [4-7], accumulation around repulsive obstacles [8], spontaneous velocity alignment [9], and active turbulence [10,11]. Interestingly, socalled linear swimmer models such as active Brownian particles (ABPs) [12-17], active Ornstein-Uhlenbeck particles [18], and run-and-tumble particles [19,20] have already been remarkably successful in theoretically describing a significant number of these nonequilibrium features. Members of this class of particles are typically endowed with a constant (average) self-propulsion whose direction changes randomly via some form of rotational diffusion (often thermal fluctuations). However, due to, for instance, an asymmetric shape [21–23], mass distribution [24], or self-propulsion mechanism [25,26], active particles also frequently self-rotate which is not included in the aforementioned models. This leads to chiral-symmetry breaking of the corresponding active motion and, at small enough densities, circular (2D) or helical (3D) trajectories. A collection of these spinning particles is usually referred to as an active chiral fluid and has been shown to exhibit many interesting collective phenomena in both simulations and experiments [23,27–39]. Understanding the influence of chirality on active matter is therefore enjoying growing

attention [40,41], but at the same time requires more involved modeling efforts to fully comprehend.

Initial chiral active matter studies have focused primarily on the low to moderate density regime [21,26,42–44], but interest is now increasingly shifting toward high densities. This has already vielded several seminal works in the context of so-called chiral crystals [27,32,45–49]. At the same time, their disordered counterpart, i.e., an active chiral glass, has received little attention. This might be attributed to the standard assumption of interactions impeding any form of active motion in the high density or glassy regime. As a result, the specific details of the active motion, whether chiral or nonchiral, should become of lesser importance upon approaching dynamical arrest. In this work we demonstrate that this naive picture is too simplistic and that chiral active motion can certainly influence glassy dynamics in highly surprising ways. We, for the first time, delve into the unique physics that emerges when a chiral fluid ventures into the glassy regime. Most notably, we introduce a new "hammering" mechanism (see Fig. 1), unique to rapidly spinning particles in high-density conditions, that can fluidize a chiral active solid.

In short, we explore the dynamics of interacting chiral active Brownian particles (CABPs) [28,42] and show that when pushed to glassy conditions our chiral fluid exhibits highly nontrivial dynamics, particularly compared to standard linear active glassy matter (that is, conventional ABPs), which has already been extensively studied in theory [50–59] and simulation [60–72]. Despite the added complexity, we are still able to present a full rationalization for all identified dynamical regimes, including the emergence of a complex

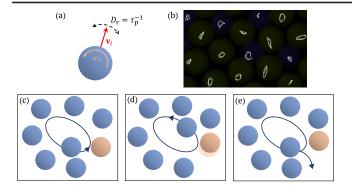


FIG. 1. (a) Visualization of a chiral active Brownian particle (CABP). (b) Example short-time trajectories (total time is equal to three spinning periods) of CABPs at large spinning frequency and persistence exhibiting the hammering effect by undergoing circular motion inside their cage of surrounding particles. (c)–(e) Schematic depiction of the hammering effect. (c),(d) For large enough persistence and spinning frequency, particles undergo back-and-forth motion inside their cage and systematically collide with the same particle whose motion is slightly altered by the collision. (e) After repeated collisions the cage of a particle is sufficiently remodeled such that the particle can break out and migrate through the material.

reentrant behavior which we explain by invoking the aforementioned hammering mechanism.

Simulation details.—As our model chiral fluid we consider a two-dimensional (2D) Kob-Andersen mixture which consists of  $N_A=650$  and  $N_B=350$  self-propelling quasihard disks of types A and B, respectively. We assume that the self-propulsion dominates over thermal fluctuations so that we can neglect passive diffusion and the equation of motion for the position  $\mathbf{r}_i$  of each particle i is given by [53]

$$\dot{\mathbf{r}}_i = \zeta^{-1} \mathbf{F}_i + \mathbf{v}_i. \tag{1}$$

Here,  $\zeta$  represents the friction constant and  $\mathbf{v}_i$  the selfpropulsion velocity acting on particle i. The interaction force  $\mathbf{F}_i = -\sum_{j \neq i} \nabla_i V_{\alpha\beta}(r_{ij})$  is obtained from a quasihard sphere power law potential  $V_{\alpha\beta}(r)=4\epsilon_{\alpha\beta}(\sigma_{\alpha\beta}/r)^{36}$  [73,74] and the interaction parameters, i.e.,  $\epsilon_{AA} = 1$ ,  $\epsilon_{AB} = 1.5$ ,  $\epsilon_{BB}=0.5,\ \sigma_{AA}=1,\ \sigma_{AB}=0.8,\ \sigma_{BB}=0.88,\ {\rm are,\ in\ com-}$ bination with setting  $\zeta = 1$ , chosen to frustrate crystallization and allow for glassy behavior [75,76]. The choice of parameters also implies that we use reduced units where  $\sigma_{AA}$ ,  $\epsilon_{AA}$ ,  $\epsilon_{AA}/k_B$ , and  $\zeta \sigma_{AA}^2/\epsilon_{AA}$  represent the units of length, energy, temperature, and time, respectively [77]. For the self-propulsion of each particle we employ the CABP scheme [28,42]. That is, the magnitude of the selfpropulsion or active speed  $v_0$  is assumed to remain constant in time t so that  $\mathbf{v}_i = v_0 \mathbf{e}_i = v_0 [\cos(\theta_i), \sin(\theta_i)]$ , while the orientation angle of the active velocity  $\theta_i$  evolves in time according to

$$\dot{\theta}_i = \chi_i + \omega_s, \tag{2}$$

with  $\omega_s$  a constant spinning frequency,  $\chi_i$  a Gaussian noise process with zero mean and variance  $\langle \chi_i(t)\chi_j(t')\rangle_{\rm noise}=2D_r\delta_{ij}\delta(t-t')$ , and  $D_r$  the rotational diffusion coefficient. As our control parameters we take  $\omega_s$ , the persistence time  $\tau_p=D_r^{-1}$ , and a so-called spinning temperature  $T_{\omega_s}=v_0^2/2\omega_s$  which represents (up to a prefactor  $4\pi\zeta$ ) a measure for the amount of energy that is dissipated by a single CABP during one circle motion.

Simulations are performed by solving the overdamped equation of motion [Eq. (1)] via a forward Euler scheme using LAMMPS [78]. We set the cutoff radius at  $r_c=2.5\sigma_{\alpha\beta}$ , fix the size of the periodic square simulation box to ensure that the number density equals  $\rho=1.2$ , run the system sufficiently long (typically between 500 and 10 000 time units) to prevent aging, and afterward track the particles over time for at least twice the initialization time. To correct for diffusive center-of-mass motion all particle positions are retrieved relative to the momentary center of mass [77].

Nonmonotonic dynamics.—We are primarily interested in characterizing how the interplay between rotational diffusion and spinning motion influences the active glassy dynamics. Therefore, we have calculated the longtime diffusion coefficient  $D = \lim_{t\to\infty} \langle \Delta \mathbf{r}_i^2(t) \rangle/4t$  of our chiral fluid for several set spinning frequencies  $\omega_s = 10, 100, 200$  (keeping a fixed value  $T_{\omega_s} = 4$  to ensure moderately supercooled behavior), while varying the persistence time. The results are plotted as a function of  $\omega_s \tau_p$  in Fig. 2 and show remarkably rich dynamics. In particular, we find initial nonmonotonic behavior with a maximum at  $\omega_s \tau_p \sim 1$ . This is followed by a form of reentrant behavior which becomes much more pronounced for higher spinning frequencies. For example, at  $\omega_s = 200$  the diffusivity reaches a minimum with  $D \sim 10^{-4}$ , which is practically

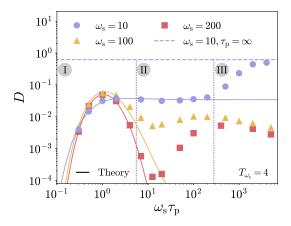


FIG. 2. The longtime diffusion coefficient D as a function of the normalized persistence time  $\omega_s \tau_p$  for several set values of  $\omega_s$  keeping  $T_{\omega_s} = 4$  fixed. The resulting dynamics show highly nontrivial behavior that can be characterized by a nonmonotonic (I), reentrant (II), and large persistence (III) regime. The dashed line indicates the infinite persistence limit (from simulations) which is only nonzero for  $\omega_s = 10$ .

a frozen system like a glass, that is seen to increase with orders of magnitude. Finally, in the limit of large persistence different asymptotic values ranging from significantly enhanced to zero dynamics are reached. To contrast these complex dynamics, we emphasize that a glassy liquid of standard ABPs at constant active speed  $v_0$  would only show a monotonic enhancement of the dynamics for increasing persistence time [50,61]. Thus, at large densities chirality has a highly nontrivial impact on active particle motion.

Moreover, we have verified that the same qualitative behavior is observed for both a different model glass former and a different set of parameters where we have fixed either the active speed  $v_0$  or the spinning frequency  $\omega_s$  instead of the spinning temperature  $T_{\omega_s}$  (see Supplemental Material, Figs. S1 and S2 [79]. We also mention that the nontrivial change of the dynamics and in particular the reentrant behavior are equally visible in the static structure factor, the self- and collective intermediate scattering function, the non-Gaussian parameter, and the dynamical susceptibility (see Supplemental Material, Figs. S3–S6 [79]). The latter two are measures for dynamical heterogeneity. The strong nonmonotonic behavior is thus a robust feature of the entire glassy phenomenology, but for convenience we continue to focus on the diffusivity.

CABP in a harmonic trap.—Our aim now is to better understand the complex dynamics, which, for convenience, we will separate in three distinct regimes (see roman numerals in Fig. 2). We first turn our attention toward regime I. Here, the persistence of particles is still relatively weak and we therefore expect that especially in this regime the local environment of particles (or their cage) acts primarily as an effective confining potential. This then motivates a comparison of our simulation results to those of a single CABP in a harmonic trap of the form  $U(r) = \kappa r^2/2$ , with r the radial distance inside the trap and  $\kappa$  its strength. In particular, we have analytically derived an expression for the longtime limit of the mean-squared displacement (MSD) of such a trapped particle (see Supplemental Material [79] for details). This yields

$$\delta \equiv \lim_{t \to \infty} \langle \Delta \mathbf{r}^2(t) \rangle = \frac{v_0^2 (1 + \frac{\omega_k}{\omega_s \tau_p})}{k^2 [\omega_k^2 + (1 + \frac{\omega_k}{\omega_s \tau_p})^2]}, \quad (3)$$

where  $k = \kappa/\zeta$ , and we have introduced the dimensionless spinning frequencies  $\omega_k = \omega_s/k$  and  $\omega_s \tau_p$ . Because of the high-density (or glassy) conditions, we can then postulate that a particle escapes its trap when it reaches a distance equal to its diameter, that is, when  $r = \sigma_{AA}$ . Assuming a Kramers-like process [84,85], the corresponding average escape time is given by

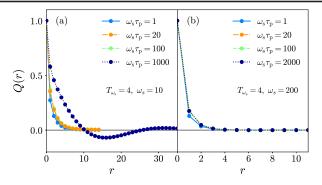
$$t_{\rm esc} = t_0 e^{\frac{\mathrm{U}(\sigma_{AA})}{\langle U \rangle}} = t_0 e^{\frac{\sigma_{AA}^2}{\delta}},\tag{4}$$

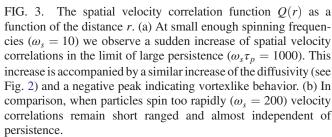
with  $t_0$  a constant prefactor and we have used that the average potential energy of the particle (which serves as our effective temperature) is equal to  $\langle U \rangle = \kappa \delta/2$  [85]. Moreover, we assume that after each escape the particle falls into a new trap with the same properties. In other words, the particle diffuses through space by hopping from trap to trap (or equivalently cage to cage). This allows us to quantitatively estimate the longtime diffusion coefficient as  $D \approx \sigma_{AA}^2/4t_{\rm esc}$ , which can be compared to our simulation results. Note that the qualitative behavior of our theoretically predicted D is thus fully determined by a single fit parameter  $t_0$ , while the absolute scale is set by the other fit parameter  $t_0$ .

The resulting theoretical predictions (fitted on the first five data points) are shown as straight lines in Fig. 2. Remarkably, we find almost quantitative agreement in regime I and approximately the same fit value of  $k \sim 10$  for all our settings (the latter is consistent with the fact that we do not change the density or interaction potential which supposedly determine this factor). This demonstrates that the interplay between rotational diffusion and spinning motion in the small persistence regime are well captured by our simple single particle model.

Collective motion.—Inspired by previous work in the literature [29] we now proceed to regime III. Here, we observe a sudden increase of the diffusivity toward the infinite persistence limit for relatively small spinning frequencies ( $\omega_s = 10$ ). For larger spinning frequencies, we instead see D decreasing and probably moving toward the infinite persistence limit of D = 0 (see Fig. S7 [79]). Thus, for  $\tau_p \to \infty$  there exists a transition from a so-called active (D > 0) to an absorbing (D = 0) state upon increasing  $\omega_s$ . This behavior is fully consistent with previous work conducted at lower densities [29]. Our work shows that this phenomenology is retained in the high density or glassy regime.

To explain why at small enough  $\omega_s$ , that is,  $\omega_s = 10$ , the diffusivity increases significantly in regime III, we employ a spatial velocity correlation function Q(r) (see Ref. [17] for a precise definition). This function measures how correlated the velocities  $\dot{\mathbf{r}}_i$  of different particles i are over a distance r. It thus serves as a proxy for local velocity alignment and cooperative motion [Q(r) = 1, 0, -1] for perfect velocity, no velocity, and antivelocity alignment, respectively]. We have plotted Q(r) for  $\omega_s = 10$  and several values of  $\omega_s \tau_p$  in Fig. 3(a). In almost all cases we see a similar rapid decay to zero implying that there exists little velocity alignment and cooperative motion is absent. However, at exactly the same point where the diffusivity has increased in regime III, i.e.,  $\omega_s \tau_p = 1000$ , we find that the decay of Q(r) suddenly becomes much more long-ranged and even develops a negative peak (both features have been checked for finite-size effects). We interpret this as the particle motion becoming more collective and vortexlike (see Fig. S8 [79] for representative





velocity fields) which explains why its overall diffusion is enhanced. Moreover, note that this collective motion is only able to emerge when the amount of rotational diffusion is small enough, that is, when we are at a large enough  $\tau_p$ .

In comparison, for  $\omega_s = 200$  we find almost no spatial velocity correlations for any value of  $\omega_s \tau_p$  [see Fig. 3(b)], which implies that no cooperative motion takes place. We expect that this is caused by particles spinning too rapidly, which prevents them from inducing any collective motion, even in the absence of rotational diffusion  $(\tau_p \to \infty)$ . Ultimately, this should lead to particles becoming trapped in circular or elliptical trajectories inside their cage, thus explaining why D goes to zero.

Hammering dynamics.—We finalize our discussion by considering the intermediate persistence regime II. Interestingly, in this regime the agreement between theory and simulation only remains intact for a relatively small spinning frequency ( $\omega_s = 10$ , see Fig. 2). In comparison, for larger spinning frequencies a competing mechanism emerges which is able to increase the longtime diffusivity from an almost glassy or dynamically arrested state  $(D \sim 10^{-4})$  with multiple orders of magnitude. We have checked that this reentrant behavior becomes even more extreme for larger values of  $\omega_s$ . The key question, therefore, is what causes the dynamics to be dramatically faster if it is not cooperative diffusion [since there are almost no spatial velocity correlations, see Fig. 3(b)]. To answer this we propose a new hammering mechanism that is distinct for rapidly spinning chiral particles at large densities [see Figs. 1(c)–1(e) for a schematic depiction]. In short, for large enough persistence and spinning frequency, particles are expected to undergo long periods of uninterrupted back-and-forth motion inside their cage. During this, they

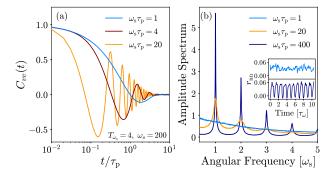


FIG. 4. (a) The velocity autocorrelation function  $C_{vv}(t)$  as a function of time t at  $\omega_s = 200$ . As  $\omega_s \tau_p > 1$  we see the emergence of oscillations which correspond to more circular particle motion inside the cage. (b) Amplitude spectrum of the nearest-neighbor distance  $r_{nn}(t)$  for several values of  $\omega_s \tau_p$ . Starting from  $\omega_s \tau_p \sim 20$ , i.e., the minimum in Fig. 2, the spectrum begins to peak at  $\omega_s$  (and its multiples). This corresponds to particles periodically colliding with the same particle inside their cage. The appearance of this hammering effect is especially visible in the inset where  $r_{nn}(t)$  for two example particles at  $\omega_s \tau_p = 1$  and 400 are plotted.

systematically collide with the same particle whose motion is slightly altered by each collision. After repeated collisions the cage of a particle is sufficiently remodeled such that the particle can break out and migrate through the material which should lead to faster dynamics.

In order to verify and explain the mechanism in more detail, we will now exclusively focus on the data obtained for  $\omega_s = 200$ ,  $T_{\omega_s} = 4$  (the red squares in Fig. 2), where the hammering effect is strongest. We start by introducing the spinning radius  $R=v_0/\omega_s$  which in this case is smaller than a particle radius, i.e., R = 0.2. If the persistence time is then also larger than the spinning period  $\tau_{\omega} = 2\pi/\omega_s$  $(\tau_p > \tau_\omega \text{ or } \omega_s \tau_p > 2\pi)$ , particles should be able to undergo full circular or elliptical motion inside their cage. This can be clearly seen when inspecting multiple short-time particle trajectories [see Fig. 1(b) and Supplemental Material, Fig. S9 [79]]. To also quantify the periodic motion, we have extracted the normalized velocity autocorrelation function  $C_{vv}(t) = \langle \dot{\mathbf{r}}_i(0) \cdot \dot{\mathbf{r}}_i(t) \rangle / \langle \dot{\mathbf{r}}_i^2 \rangle$  for a subset of  $\omega_s \tau_p$  values and plotted them in Fig. 4(a). In accordance with the more circular trajectories, we observe the emergence of oscillations, which roughly start when  $\omega_s \tau_p \gtrsim 2\pi$  and become longer lived for increasing persistence.

Having established that particles increasingly go in circles inside their cage, we also want to show that this is indeed accompanied by systematic collisions with one (or multiple) of its surrounding particles. For this reason we have, based on short-time trajectories, extracted the nearest neighbor distance for a given particle i,  $r_{\rm nn}(t) = \min\{|{\bf r}_i(t)-{\bf r}_j(t)|\}_{j\neq i}$ , as a function of time. Two of these have been plotted as an inset in Fig. 4(b) and confirm the notion of repeated collisions. Specifically, for  $\omega_s \tau_p = 1$ , i.e.,  $\tau_p < \tau_\omega$ , the nearest neighbor distance appears to be

completely random, that is, the particle undergoes random collisions with all neighboring particles. For  $\omega_s \tau_p = 400$ , i.e.,  $\tau_p \gg \tau_\omega$ , the nearest neighbor distance is instead very periodic indicating that the particle collides, moves away, collides again, and so on. We have verified that these collisions are with the same neighboring particle. To check that this behavior occurs throughout our material we have also calculated the spectrum of  $r_{\rm nn}(t)$  for all particles [see Fig. 4(b)]. We see that, starting from  $\omega_s \tau_p \sim 20$  (the minimum in Fig. 2) the spectrum begins to peak at  $\omega_s$  (and its multiples), thus further corroborating the idea of particles periodically colliding.

Overall, these results show that a hammering mechanism is indeed present in our active chiral fluid and is likely to be responsible for the observed enhanced dynamics in regime II. Moreover, it also explains why the minimum of the dynamics is roughly at  $\tau_p \sim 2\tau_\omega$  since only from this point onward are particles able to, on average, make multiple systematic collisions with the same neighboring particle and start capitalizing on the hammering effect.

Conclusion.—To conclude, our Letter demonstrates that chiral glassy fluids exhibit a remarkably rich dynamical phenomenology (especially when contrasted with their nonchiral counterpart [50,61]) that can be characterized by a nonmonotonic (I), reentrant (II), and large persistence (III) regime. We have shown how the initial behavior (I) is fully explained by treating the surroundings of a particle as a harmonic trap and considering cage hopping of a single CABP between such traps. In the limit of extremely weak rotational diffusion (III), we have observed either speeding up or slowing down of the dynamics which is related to the (in)ability of particles to align their respective velocities and induce collective swirlinglike motion. Finally, to rationalize the surprising (but for large spinning frequencies highly significant) reentrant behavior (II), we have introduced and demonstrated the existence of a new hammering mechanism that is distinct for rapidly spinning particles at high densities. Overall, our results pay testimony to the fact that chirality (already in its simplest form) gives rise to a plethora of nontrivial behavior, even in the glassy limit where interactions usually dominate dynamics. It would be interesting to see whether these regimes and specifically the hammering mechanism can also be observed for so-called spinners where active motion is induced via rotational couplings [35] or if it can possibly be exploited in an experimental setting involving, for instance, active granular rotors or colloids. For the latter one might also have to consider the role of translational diffusion, which could hinder collective motion or disrupt circular motion inside the cage. Alternatively, one can think of chiral active probe particles possibly utilizing the hammering effect to help extract material properties or navigate through a soft dense environment such as gels [86–90].

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- [1] C. Bechinger, R. Di Leonardo, H. Löwen, C. Reichhardt, G. Volpe, and G. Volpe, Rev. Mod. Phys. 88, 045006 (2016).
- [2] S. Ramaswamy, Annu. Rev. Condens. Matter Phys. 1, 323 (2010).
- [3] M. C. Marchetti, J. F. Joanny, S. Ramaswamy, T. B. Liverpool, J. Prost, M. Rao, and R. A. Simha, Rev. Mod. Phys. 85, 1143 (2013).
- [4] I. Buttinoni, J. Bialké, F. Kümmel, H. Löwen, C. Bechinger, and T. Speck, Phys. Rev. Lett. 110, 238301 (2013).
- [5] F. Ginot, I. Theurkauff, F. Detcheverry, C. Ybert, and C. Cottin-Bizonne, Nat. Commun. 9, 696 (2018).
- [6] J. Palacci, S. Sacanna, A. P. Steinberg, D. J. Pine, and P. M. Chaikin, Science 339, 936 (2013).
- [7] M. N. van der Linden, L. C. Alexander, D. G. A. L. Aarts, and O. Dauchot, Phys. Rev. Lett. 123, 098001 (2019).
- [8] A. P. Berke, L. Turner, H. C. Berg, and E. Lauga, Phys. Rev. Lett. **101**, 038102 (2008).
- [9] J. Deseigne, O. Dauchot, and H. Chaté, Phys. Rev. Lett. 105, 098001 (2010).
- [10] L. Giomi, Phys. Rev. X 5, 031003 (2015).
- [11] R. Alert, J. Casademunt, and J.-F. Joanny, Annu. Rev. Condens. Matter Phys. 13, 143 (2022).
- [12] Y. Fily and M. C. Marchetti, Phys. Rev. Lett. 108, 235702 (2012).
- [13] J. Bialké, T. Speck, and H. Löwen, Phys. Rev. Lett. 108, 168301 (2012).
- [14] G. S. Redner, M. F. Hagan, and A. Baskaran, Phys. Rev. Lett. 110, 055701 (2013).
- [15] C. B. Caporusso, P. Digregorio, D. Levis, L. F. Cugliandolo, and G. Gonnella, Phys. Rev. Lett. 125, 178004 (2020).
- [16] R. Ni, M. A. Cohen Stuart, and P. G. Bolhuis, Phys. Rev. Lett. 114, 018302 (2015).
- [17] L. Caprini, U. Marini Bettolo Marconi, and A. Puglisi, Phys. Rev. Lett. 124, 078001 (2020).
- [18] C. Maggi, U. M. B. Marconi, N. Gnan, and R. Di Leonardo, Sci. Rep. 5, 10742 (2015).
- [19] J. Tailleur and M. E. Cates, Phys. Rev. Lett. 100, 218103 (2008).
- [20] M. E. Cates and J. Tailleur, Europhys. Lett. **101**, 20010 (2013).
- [21] F. Kümmel, B. ten Hagen, R. Wittkowski, I. Buttinoni, R. Eichhorn, G. Volpe, H. Löwen, and C. Bechinger, Phys. Rev. Lett. 110, 198302 (2013).
- [22] P. Patra, K. Beyer, A. Jaiswal, A. Battista, K. Rohr, F. Frischknecht, and U. S. Schwarz, Nat. Phys. 18, 586 (2022).
- [23] B. Zhang, A. Sokolov, and A. Snezhko, Nat. Commun. 11, 4401 (2020).
- [24] A. I. Campbell, R. Wittkowski, B. ten Hagen, H. Löwen, and S. J. Ebbens, J. Chem. Phys. 147, 084905 (2017).
- [25] R. J. Archer, A. I. Campbell, and S. J. Ebbens, Soft Matter 11, 6872 (2015).

- [26] E. Lauga, W. R. DiLuzio, G. M. Whitesides, and H. A. Stone, Biophys. J. 90, 400 (2006).
- [27] Z.-F. Huang, A. M. Menzel, and H. Löwen, Phys. Rev. Lett. 125, 218002 (2020).
- [28] Z. Ma and R. Ni, J. Chem. Phys. 156, 021102 (2022).
- [29] Q.-L. Lei, M. P. Ciamarra, and R. Ni, Sci. Adv. 5, eaau7423 (2019).
- [30] B. Liebchen and D. Levis, Phys. Rev. Lett. 119, 058002 (2017).
- [31] C. Chen, S. Liu, X.-q. Shi, H. Chaté, and Y. Wu, Nature (London) 542, 210 (2017).
- [32] T. H. Tan, A. Mietke, J. Li, Y. Chen, H. Higinbotham, P. J. Foster, S. Gokhale, J. Dunkel, and N. Fakhri, Nature (London) 607, 287 (2022).
- [33] C. Scholz, A. Ldov, T. Pöschel, M. Engel, and H. Löwen, Sci. Adv. 7, eabf8998 (2021).
- [34] V. Soni, E. S. Bililign, S. Magkiriadou, S. Sacanna, D. Bartolo, M. J. Shelley, and W. T. M. Irvine, Nat. Phys. 15, 1188 (2019).
- [35] M. Han, M. Fruchart, C. Scheibner, S. Vaikuntanathan, J. J. de Pablo, and V. Vitelli, Nat. Phys. 17, 1260 (2021).
- [36] C. Reichhardt and C. J. O. Reichhardt, Phys. Rev. E 100, 012604 (2019).
- [37] G.-J. Liao and S. H. L. Klapp, Soft Matter 14, 7873 (2018).
- [38] G.-J. Liao and S. H. L. Klapp, Soft Matter 17, 6833 (2021).
- [39] M. Huang, W. Hu, S. Yang, Q.-X. Liu, and H. P. Zhang, PNAS 118, e2100493118 (2021).
- [40] M. J. Bowick, N. Fakhri, M. C. Marchetti, and S. Ramaswamy, Phys. Rev. X 12, 010501 (2022).
- [41] B. Liebchen and D. Levis, Europhys. Lett. **139**, 67001 (2022).
- [42] S. van Teeffelen and H. Löwen, Phys. Rev. E **78**, 020101(R) (2008).
- [43] T. Ohta and T. Ohkuma, Phys. Rev. Lett. **102**, 154101 (2009).
- [44] R. Di Leonardo, D. Dell'Arciprete, L. Angelani, and V. Iebba, Phys. Rev. Lett. 106, 038101 (2011).
- [45] A. P. Petroff, X.-L. Wu, and A. Libchaber, Phys. Rev. Lett. 114, 158102 (2015).
- [46] J. Yan, S. C. Bae, and S. Granick, Soft Matter 11, 147 (2015).
- [47] K. Drescher, K. C. Leptos, I. Tuval, T. Ishikawa, T. J. Pedley, and R. E. Goldstein, Phys. Rev. Lett. 102, 168101 (2009).
- [48] K. Drescher, R. E. Goldstein, N. Michel, M. Polin, and I. Tuval, Phys. Rev. Lett. 105, 168101 (2010).
- [49] T. Ishikawa, T. J. Pedley, K. Drescher, and R. E. Goldstein, J. Fluid. Mech. 903, A11 (2020).
- [50] A. Liluashvili, J. Ónody, and T. Voigtmann, Phys. Rev. E 96, 062608 (2017).
- [51] G. Szamel, J. Chem. Phys. 150, 124901 (2019).
- [52] G. Szamel, E. Flenner, and L. Berthier, Phys. Rev. E 91, 062304 (2015).
- [53] G. Szamel, Phys. Rev. E 93, 012603 (2016).
- [54] M. Feng and Z. Hou, Soft Matter 13, 4464 (2017).
- [55] L. Berthier and J. Kurchan, Nat. Phys. 9, 310 (2013).
- [56] J. Reichert, S. Mandal, and T. Voigtmann, Phys. Rev. E 104, 044608 (2021).

- [57] J. Reichert and T. Voigtmann, Soft Matter 17, 10492 (2021).
- [58] J. Reichert, L. F. Granz, and T. Voigtmann, Eur. Phys. J. E 44, 27 (2021).
- [59] S. K. Nandi, R. Mandal, P. J. Bhuyan, C. Dasgupta, M. Rao, and N. S. Gov, Proc. Natl. Acad. Sci. U.S.A. 115, 7688 (2018).
- [60] L. Berthier, Phys. Rev. Lett. 112, 220602 (2014).
- [61] R. Ni, M. A. C. Stuart, and M. Dijkstra, Nat. Commun. 4, 2704 (2013).
- [62] L. Berthier, E. Flenner, and G. Szamel, New J. Phys. 19, 125006 (2017).
- [63] E. Flenner and G. Szamel, Phys. Rev. E 102, 022607 (2020).
- [64] E. Flenner, G. Szamel, and L. Berthier, Soft Matter 12, 7136 (2016).
- [65] S. Henkes, Y. Fily, and M. C. Marchetti, Phys. Rev. E 84, 040301(R) (2011).
- [66] R. Mandal and P. Sollich, Phys. Rev. Lett. 125, 218001 (2020).
- [67] G. Janzen and L. M. C. Janssen, Phys. Rev. Res. 4, L012038 (2022).
- [68] D. Bi, X. Yang, M. C. Marchetti, and M. L. Manning, Phys. Rev. X 6, 021011 (2016).
- [69] M. Paoluzzi, D. Levis, and I. Pagonabarraga, Commun. Phys. 5, 111 (2022).
- [70] V. E. Debets, X. M. de Wit, and L. M. C. Janssen, Phys. Rev. Lett. 127, 278002 (2021).
- [71] Y.-E. Keta, R. L. Jack, and L. Berthier, Phys. Rev. Lett. 129, 048002 (2022).
- [72] V. E. Debets and L. M. C. Janssen, Phys. Rev. Res. 4, L042033 (2022).
- [73] F. Weysser, A. M. Puertas, M. Fuchs, and T. Voigtmann, Phys. Rev. E 82, 011504 (2010).
- [74] E. Lange, J. B. Caballero, A. M. Puertas, and M. Fuchs, J. Chem. Phys. 130, 174903 (2009).
- [75] W. Kob and H. C. Andersen, Phys. Rev. Lett. **73**, 1376 (1994).
- [76] C. D. Michele, F. Sciortino, and A. Coniglio, J. Phys. Condens. Matter 16, L489 (2004).
- [77] E. Flenner and G. Szamel, Phys. Rev. E 72, 031508 (2005).
- [78] S. Plimpton, J. Comput. Phys. 117, 1 (1995).
- [79] See Supplemental Material at <a href="http://link.aps.org/supplemental/10.1103/PhysRevLett.130.058201">http://link.aps.org/supplemental/10.1103/PhysRevLett.130.058201</a> for a detailed derivation of the MSD of a CABP in a harmonic trap, example flow fields, and additional data for different settings and of more involved glassy features, which includes Refs. [80–83].
- [80] D. Chaudhuri and A. Dhar, J. Stat. Mech. (2021) 013207.
- [81] H. Löwen, J. Chem. Phys. 152, 040901 (2020).
- [82] S. Jahanshahi, H. Löwen, and B. ten Hagen, Phys. Rev. E **95**, 022606 (2017).
- [83] D. N. Perera and P. Harrowell, Phys. Rev. Lett. 81, 120 (1998).
- [84] C. W. Gardiner, Handbook of Stochastic Methods (Springer, New York, 1985), Vol. 3.
- [85] E. Woillez, Y. Kafri, and N. S. Gov, Phys. Rev. Lett. 124, 118002 (2020).
- [86] C. Kurzthaler, S. Mandal, T. Bhattacharjee, H. Löwen, S. S. Datta, and H. A. Stone, Nat. Commun. 12, 7088 (2021).

- [87] C. Abaurrea-Velasco, C. Lozano, C. Bechinger, and J. de Graaf, Phys. Rev. Lett. **125**, 258002 (2020).
- [88] J. R. Gomez-Solano, A. Blokhuis, and C. Bechinger, Phys. Rev. Lett. **116**, 138301 (2016).
- [89] C. Lozano, J. R. Gomez-Solano, and C. Bechinger, Nat. Mater. 18, 1118 (2019).
- [90] E. Irani, Z. Mokhtari, and A. Zippelius, Phys. Rev. Lett. **128**, 144501 (2022).